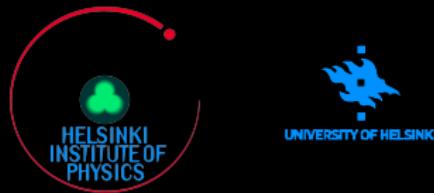


Backreaction and FLRW consistency conditions

Francesco MONTANARI



FM, Syksy Räsänen [arXiv:1709.06022]

LAM Marseille – 2018, May 31

FLRW consistency condition

Clarkson, Basset, Hui-Ching Lu [arxiv:0712.3457]

$$d(z) = \frac{1}{\sqrt{k}} \sin \left(\sqrt{k} \int_0^z \frac{d\tilde{z}}{h(\tilde{z})} \right) \quad \Rightarrow \quad \boxed{k = \frac{1 - h^2 d'^2}{d^2} \equiv k_H(z)}$$

$$d(z) = H_0(1+z)D_A(z)$$

$$h(z) = H(z)/H_0$$

Backreaction

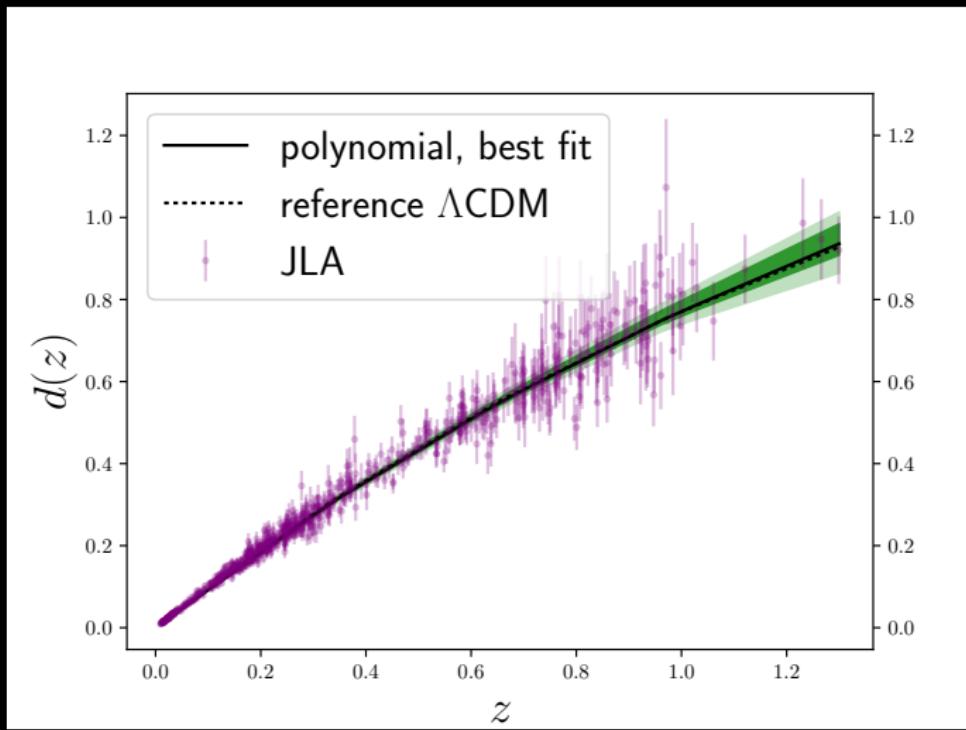
Sachs optical equation

$$h \frac{d}{dz} \left[(1+z)^2 h d'_A \right] = -\frac{3}{2} \Omega_{m0} (1+z)^3 d_A$$

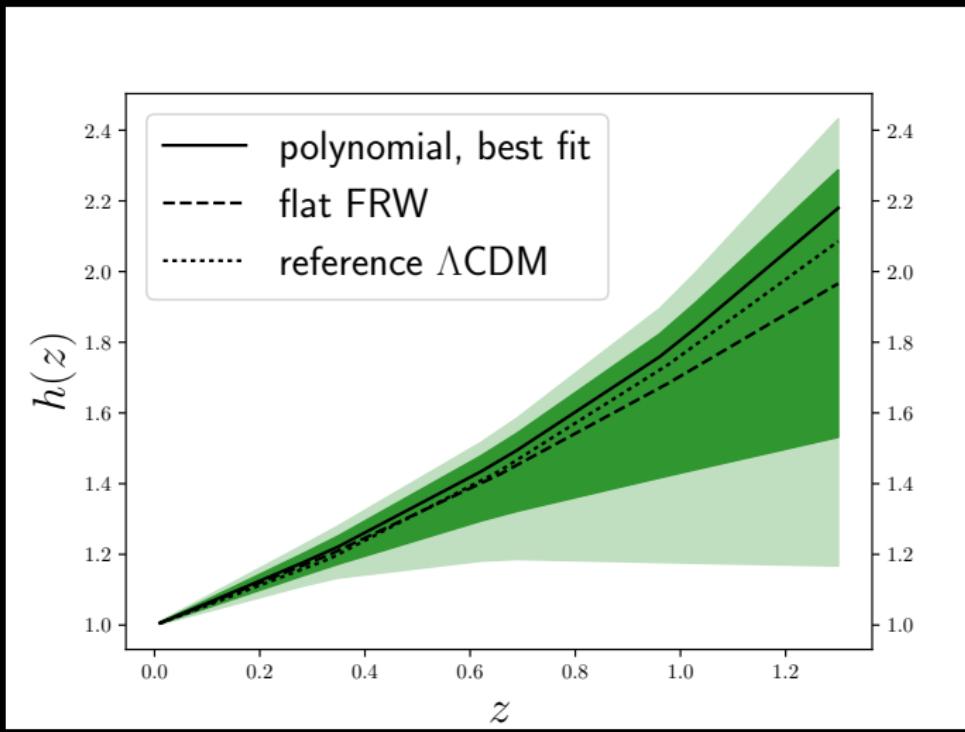
gives

$$h(z)^2 = \frac{1}{(1+z)^4 (d'_A)^2} \left[1 - 3\Omega_{m0} \int_0^z d\tilde{z} (1+\tilde{z})^5 d_A(\tilde{z}) d'_A(\tilde{z}) \right]$$

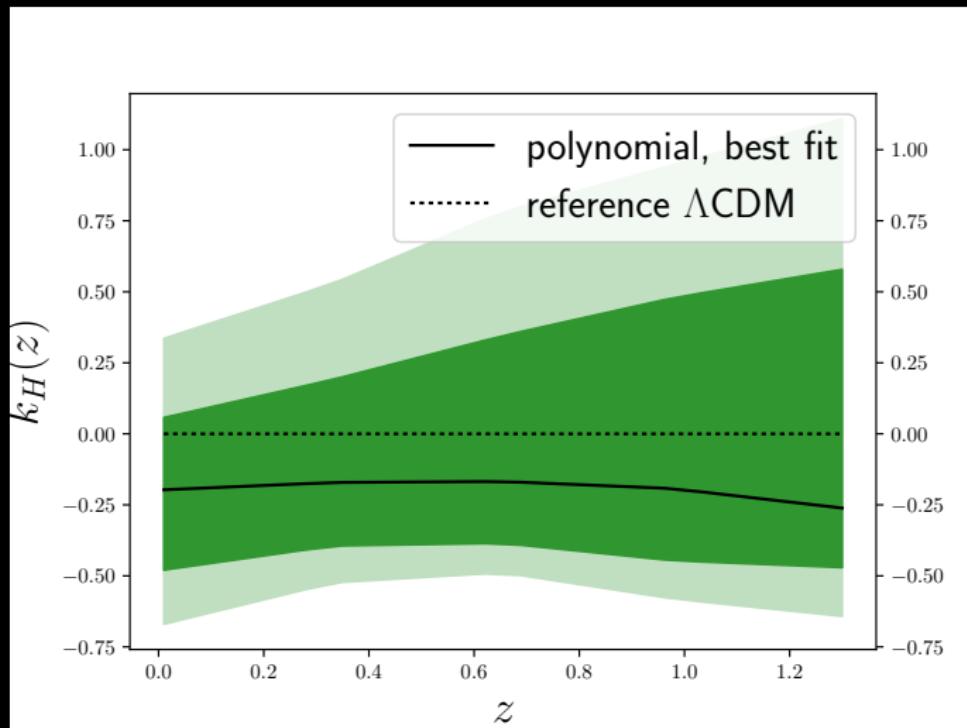
Distance



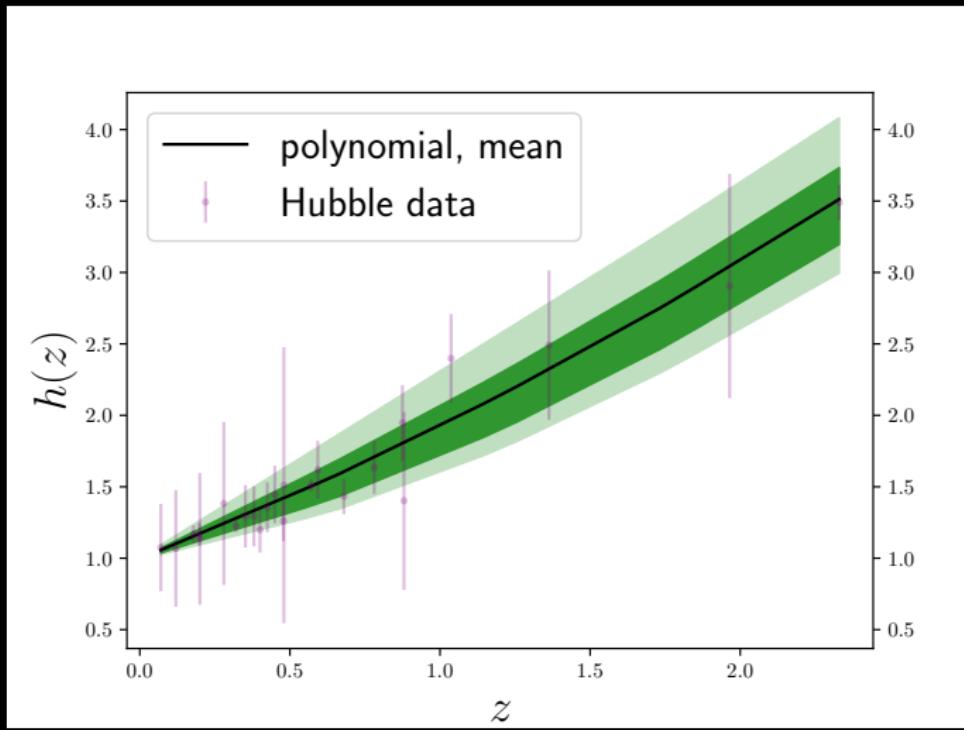
Backreaction expansion rate



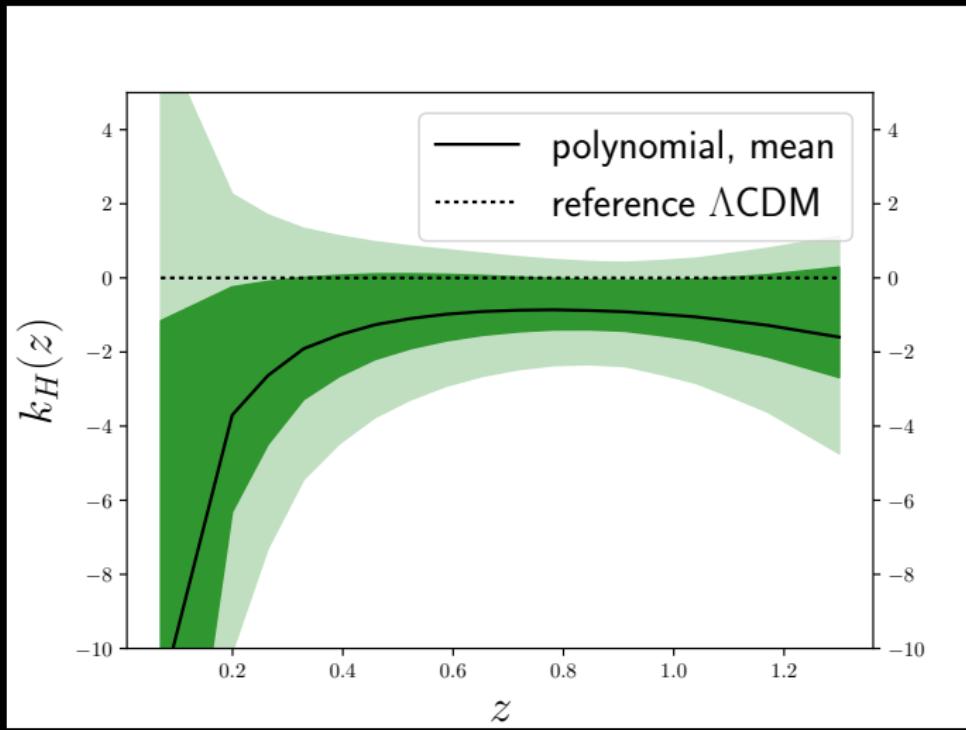
Backreaction prediction



Expansion rate from BC03 cosmic clocks + BAO



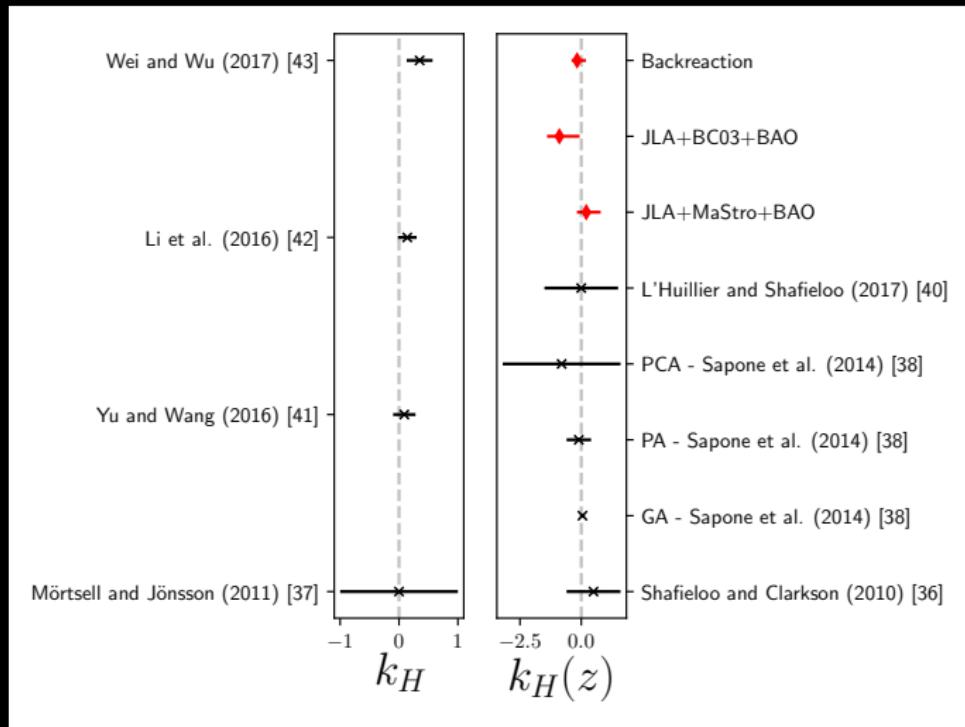
FLRW null test



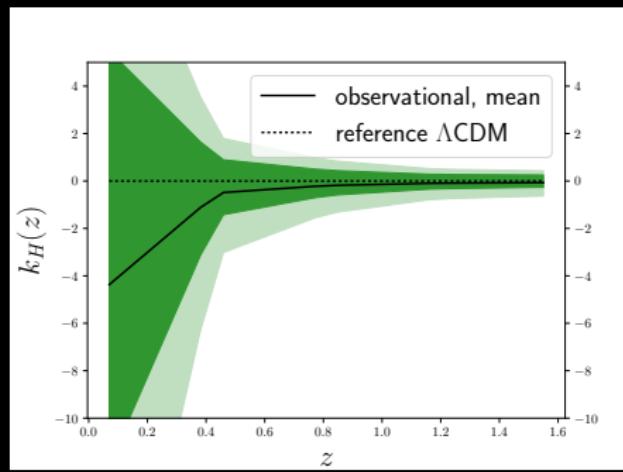
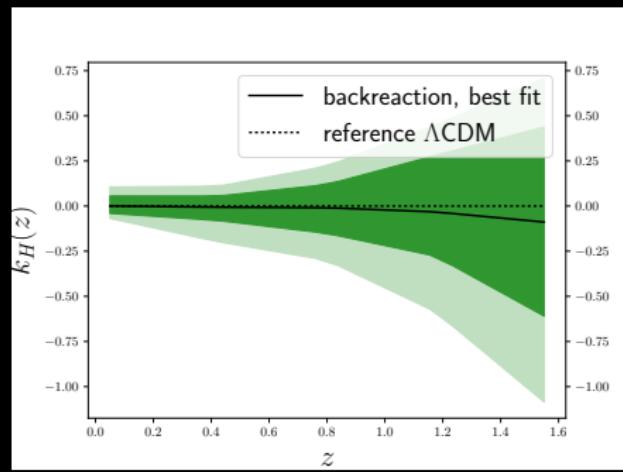
Hubble parameter today

| H_0 [km/Mpc/s] | BC03 | BC03+BAO | MaStro | MaStro+BAO |
|------------------|----------------------|----------------------|----------------------|----------------------|
| polynomial | $66.8^{+6.1}_{-6.3}$ | $64.2^{+5.2}_{-3.9}$ | 70.7^{+12}_{-13} | $67.7^{+4.9}_{-4.8}$ |
| spline | $68.8^{+7.3}_{-7.1}$ | $62.5^{+4.6}_{-4.6}$ | 69.0^{+15}_{-16} | $68.7^{+5}_{-5.3}$ |
| Λ CDM | $68.4^{+6.2}_{-6.3}$ | $61.7^{+4.5}_{-4.5}$ | $79.6^{+6.7}_{-7.5}$ | $67.7^{+5.3}_{-4.8}$ |

k_H constraints



Forecast LSST+Euclid



Conclusions

- Backreaction prediction (best 95% C.I.): $-0.7 < k_H < 0.4$.
- Geometric, model independent test (best 95% C.I.):
 - JLA+BC03+BAO: $-2.32 < k_H < 0.4$.
 - JLA+MaStro+BAO: $-0.86 < k_H < 1.13$.
- Non-trivial degeneracies in parameter space.
- LSST+Euclid forecast, improvement on $k_H(z)$ by factor 6 (backreaction) and 3 (observational).
- Further consistency conditions: $k_p(z)$ (parallax), $k_S(z_l, z_s)$ (distances sum rule).